

**Appendix by:**

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(not implemented in code)

**Anomalous Heating (LHDI)**

- Model Based on
- A. E. Robson, Phys. Fluid B3, 1481(1991)
- H. Bruzzone, Nucleonika 46, S3(2001)
- L. Bernal, H. Bruzzone, Plasma Phys.&Cont. Fusion, 41, 223(2002)
- Chittenden, Phys. Plasma, 2, 1242(1995).



# Low-Hybrid-Drift Instability

- It is usually originated at the later stage of implosion
- It is triggered whenever the electron drift velocity ( $v_d$ ) is greater or equal to the ion thermal speed ( $v_T$ )
- The main effect is modification of electron-ion collision frequency  $\nu_{ei}$  through additional effects (due to wave particle interaction), introduces collision frequency  $\nu^*$
- It is important in all the situations wherever the magnetic field  $B$  is perpendicular the perturbations ( $B \perp k$ ) or  $k \cdot B = 0$  and  $k_{\parallel} = 0$ .



- In order to cause electron scattering and contribute to the anomalous resistivity the wave length should be shorter or comparable to the electron gyro-radius i.e.  $\lambda_{\square}(1/k_{\square}) \leq r_L$ .
- The natural frequency of oscillation for the low-hybrid wave is

$$\omega_{LH} = \omega_{ce} \sqrt{\frac{Z_{eff} m_e}{A m_p}}$$

- In high threshold (for  $v_d = \beta v_{Ti}$  for  $\beta=(2-4)$ ), with  $T_e \gg T_i$  (even for larger effective charge  $T_e > (7T_i/Z_{eff})$ ), the Ion-acoustic instability is important.



- If the electron drift velocity  $v_d$  is very much high than the ion speed ( $\beta > 4$ ) and the ( $T_e \gg T_i$  or even  $T_i > T_e$ ), the role of two-stream or Buneman instability become important.

### *The Role of LHDI :*

The main effect of this wave-particle interaction is to enhance the effective collisional frequency (in addition to the usual electron-ion collisional frequency), causing electrical resistivity

$$\eta = \frac{m_e}{ne^2} (v_{ei} + v^*) = \eta_{\perp} + \eta^*$$

For the pinch environment (including plasma focus) with  $\omega_{pe} > \omega_{ce}$   $\{ \omega_{pe} \approx (4\pi ne/m_e)^{0.5} \sim 5.64 \times 10^4 n_e^{1/2}$  (in  $\text{cm}^{-3}$ ) and  $\omega_{ce} \approx (eB/m_e c) \sim 1.76 \times 10^7 B$  (in Gauss)  $\}$ .

The effective frequency can be written as

$$v^* = \xi \omega_{LH} \left( \frac{v_d}{v_{Ti}} \right)^2 \dots \text{where} \dots \xi = \frac{1}{2} \sqrt{\frac{\pi}{2}}$$



$$\omega_{LH} = (\omega_{ce} \omega_{ci})^{\frac{1}{2}} = \left( \frac{eB}{m_e} \frac{z_{eff} eB}{m_i} \right)^{0.5} = \omega_{ce} \left( \frac{z_{eff} m_e}{Am_p} \right)^{0.5}$$

■ Where  $m_i = Am_p$  ( $A$  is ionic mass and  $m_p$  is mass of proton),  $z_{eff}$  is the effective charge state. Then we can write

$$v^* = 0.5 \sqrt{\frac{\pi}{2}} \omega_{ce} \sqrt{\frac{z_{eff} m_e}{Am_p}} \left( \frac{v_d}{v_{Ti}} \right)^2$$

Where  $v_{Ti} = (2KT/m_i)^{0.5} = (2KT/Am_p)^{0.5}$  is the ion thermal speed,  $v_d = (J/en)$  is the electron drift velocity.

■ The anomalous heating caused by LHDI not only heat up the electrons but mostly the ions .



- The Joule heating per unit volume goes like
- (1)  $(0.13\eta^* + 0.87 \eta_{\square}) J^2$  to electron
- (2)  $(0.87\eta^* + 0.87 \eta_{\square}) J^2$  to ions
- Therefore

$$v^* = 0.5 \sqrt{\frac{\pi}{2} \left( \frac{eB}{m_e} \right)} \sqrt{\frac{Z_{eff} m_e}{m_p A}} \left( \frac{J / en}{\sqrt{\frac{2KT}{m_i}}} \right)^2$$

- Simplifying the above equation one gets

$$v^* = 0.5 \sqrt{\frac{\pi}{2} \left( \frac{42.9}{en^2} \right)} \sqrt{\frac{A}{Z_{eff}}} \left( \frac{J^2 B}{2KT} \right)$$

Or using the numerical value of constants

$$v^* = 1.68 \times 10^{20} \sqrt{\frac{A}{Z_{eff}}} \left( \frac{J^2 B}{2KT} \right) \frac{1}{n^2}$$



- Now using the value of  $v^*$ , we have

$$\eta^* = \frac{m_e v^*}{e^2 n} = \frac{m_e}{e^2 n} (1.68 \times 10^{20} \sqrt{\frac{A}{Z_{eff}} \left( \frac{J^2 B}{2KT} \right) \frac{1}{n^2}})$$

- Simplifying and putting the values of constants the above equation becomes

$$\eta^* = 5.97 \times 10^{27} \sqrt{\frac{A}{Z_{eff}} \left( \frac{J^2 B}{2KT} \right) \frac{1}{n^3}}$$

- The total resistive heating power density becomes
- $P_{ohm} = (1 + (\eta^* / \eta_{\square})) \eta_{\square} J^2$  where  $\eta_{\square} = 65.3 \times (Z \ln \Lambda / T^{(3/2)})$  is the classical Spitzer resistivity. But power per unit length  $dQ/dt = P_{ohm} \pi r^2$ .



- Therefore

$$\frac{dQ}{dt} = \left( 1 + 0.918 \times 10^{26} \frac{1}{Z \ln \Lambda} \sqrt{\frac{A}{Z_{eff}}} \left( \frac{T^{1/2} J^2 B}{2n^3} \right) \right) \frac{65.3 Z \ln \Lambda^2}{\pi r^2 T^{3/2}}$$

- Where  $J = (I/\pi r^2)$  is the current density. The power heating

$$\frac{dQ}{dt} = \left( 1 + 0.918 \times 10^{26} \frac{1}{Z \ln \Lambda} \sqrt{\frac{A}{Z_{eff}}} \left( \frac{T^{1/2} J^2 B}{2n^3} \right) \right) \frac{65.3 Z_{eff} z_f \ln \Lambda^2}{\pi r^2 T^{3/2}}$$



# Collisional Effects

- Ref. K. N. Koshelev and N.R. Pereira, JAP 69, R21(1991)
- Consider the case when the ions have two modes of excitation
- (1) Spontaneous excitation with de-excitation rate  $W_r$
- (2) Collisional de-excitation (excitation due to collision with electron) rate depends upon the density of electron ( $n_e \sim n_i \sim n$ ) i. e.  $W_e \sim n_e w_e$
- The total de-excitation rate could be  $W = W_r + n_e w_e$ . In low density coronal case the power density of the line radiation  $P_1 = n_i n_e Z_{\text{eff}} X_1$ , where  $X_1 = 4.6 \cdot 10^{-29} (Z_n/T)$  is the coefficient of the line radiation, and  $Z_n$  is the nuclear charge and  $Z_{\text{eff}}$  is the effective charge state of the ion.



- In high density (collisional case) the energy loss coefficient  $X_1'$  due line radiation is reduce by  $\frac{w_r}{(w_r + n_e w_e)} = \frac{1}{(1 + \frac{n_e w_e}{w_r})}$
- And thus

$$X_l' = X_l \left( \frac{1}{(1 + \frac{n_e w_e}{w_r})} \right)$$

- Therefore, the power loss becomes

$$P_l = n_i n_e Z_{eff} X_l \left( \frac{1}{(1 + \frac{n_e w_e}{w_r})} \right)$$

- The radiated power per unit length  $Q_1' = P_1' \pi r^2$ . For a case of dense plasma ( $n_e w_e \gg w_r$ ) the coefficient ( $w_e/w_r$ ) (for K and L-ion radiation), for analytical modeling, can be

- $\frac{w_e}{w_r} \approx 10^{-14} / T^{7/2} \dots cm^3$



- Therefore, for  $n_e w_e \gg w_r$ , the radiation power loss can be re-written as

$$\begin{aligned}
 P_l &= n_i n_e Z_{eff} \left( 4 \times 10^{-29} \frac{Z_n^4}{T} \left( \frac{1}{\frac{n_e w_e}{w_r}} \right) \right) \\
 &= n_i Z_{eff} \left( 4 \times 10^{-29} \frac{Z_n^4}{T} \frac{1}{10^{-14} / T^{7/2}} \right) \\
 &= n_i Z_{eff} \left( 4 \times 10^{-15} Z_n^4 T^{5/2} \right)
 \end{aligned}$$

in watt/cm<sup>3</sup>. The line radiation power per unit length  $Q_l$

$$Q_l = n_i Z_{eff} \left( 4 \times 10^{-15} Z_n^4 T^{5/2} \pi r^2 \right)$$



# Opacity (radiation self-absorption)

- The radiations are emitted by the plasma if its mean free-path is larger than the size of the plasma  $r$ . If the optical depth  $r/l_\nu \gg 1$  (the mean free-path  $l_\nu <$  the plasma size), the photon excited large number  $M$  ions before reaching the edge, thus  $M = r/l_\nu$ . But when  $r$  decreases the density increases ( $N = n_i \pi r^2$ ), the mean free path decreases  $l_\nu \sim (1/n_i)$ , thus  $M$  increases (as  $M \sim (1/r)$ ).
- Without collisional de-excitation these  $M$  ions re-emit photons in arbitrary direction. Multiple scattering together with collisional de-excitation deposit the photon energy back into the plasma.



- The probability that the photon is re-emitted in one cycle (including electron collisions) is

$$p = \frac{1}{\left(1 + \frac{n_e w_e}{w_r}\right)}$$

- The probability of the photon to escape from the plasma without being deposited is  $\beta = P^M$ , where M is average number of collisions per photon.
- For the case of total frequency redistribution inside the doppler line contour an acceptable approximation is  $M = \alpha(r/l_\nu)$ ; where  $\alpha$  depends upon the geometry, optical depth (as  $\ln(r/l_\nu)$ ) and line density  $N$ ; typically  $\alpha = 5$ ; so  $M = 5(r/l_\nu)$ .



- The mean free path  $l_\nu = \frac{1}{n_i \sigma_\nu}$ ; where  $\sigma_\nu$  is the cross-section of the absorption of the photon depending upon the *line width*, *photon energy*, and is proportional to the *statistical weight*  $\omega_g$  of the ground state of the ion (likely to be present in the plasma) e.g.  $\omega_g = 10$ .
- Assuming the temperature  $T = 2 h \nu$  and  $n_e = Z_{eff} n_i$ ; the mean free path becomes (in meter)

$$l_\nu = 3 \times 10^{13} \frac{Z_{eff} T^{\frac{3}{2}} (eV)}{n_e Z_n^{\frac{1}{2}}}$$

- Now the power density due to line radiation  $P_1' = P_1' p^{1+M}$  and radiation loss/unit length  $Q' = P_1' \pi r^2$ . The photonic excitation number  $M$  (in numbers) becomes

$$M = \frac{5rZ_n^{\frac{1}{2}}n}{3 \times 10^{11} Z_{eff} T^{\frac{3}{2}}} = 1.66 \times 10^{-11} \frac{rZ_n^{\frac{1}{2}}n}{Z_{eff} T^{\frac{3}{2}}}$$



- The power loss (i. )due line radiation will become

$$\frac{dQ_l}{dt} = n_i n_e Z_{eff} z_f \pi r^2 \left( 4 \times 10^{-29} \frac{Z_n^4}{T} \left( \frac{1}{1 + \frac{n_e w_e}{w_r}} \right)^{1+M} \right)$$

$$= n_i Z_{eff} z_f \pi r^2 \left( 4 \times 10^{-29} \frac{Z_n^4}{T} \left( \frac{1}{10^{-14} / T^{7/2}} \right)^{1+M} \right)$$



# Opacity and Critical Radius

- The radius of the plasma is called critical radius when the plasma radiation goes from *volume like* to *surface like*. This stage occur when a sizable fraction of photon is reabsorbed
- i.e.  $p^{1+M} = I/e = 0.37$ . Assuming that in using  $M \gg 1$  and  $n_e w_e / w_r \ll 1$  in  $p = (1/(1 + (n_e w_e / w_r)))$ , and upon expansion  $p^{1+M} \sim (M n_e w_e / w_r)$ , give critical radius (in cm)

$$r_c \approx 1.2 \times 10^5 \frac{Z_n^{1/6} I^{4/3}}{Z_{eff}^{1/2} T^{7/3}}$$

- The radiation losses are assume to be surface like for plasma aize smaller than  $r_c$  and surface-like radiation loss per unit length  $Q_s$  (Watt/cm)

$$Q_s = 2.7 \times 10^{-2} r Z n^{7/2} Z_{eff}^{1/2} T^4$$



- Surface like radiations are much weaker than the black body emission. A black body of cylinder with radius  $r$  radiates

$$Q_{BB} = 2\pi r \sigma_{SB} T^4$$

- watts per unit length, where  $\sigma_{SB} = 10^5$  (W/cm<sup>2</sup>-eV<sup>4</sup>) is Stefan Boltzman constant. The emissivity  $\varepsilon$

$$\varepsilon = Q_s / Q_{BB} \approx 4 \times 10^{-7} Z_{eff}^{7/2}$$

- and is about 1/30 for iron.